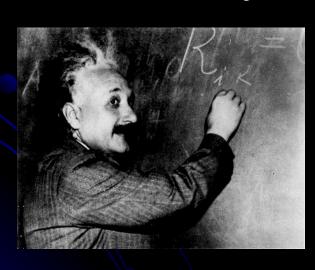
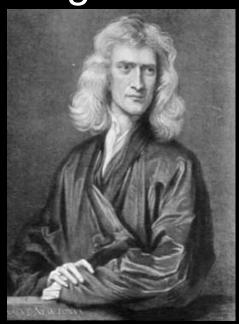
# Einstein's Tests of General Relativity through the Eyes of Newton

James Espinosa

Department of Physics

University of West Georgia



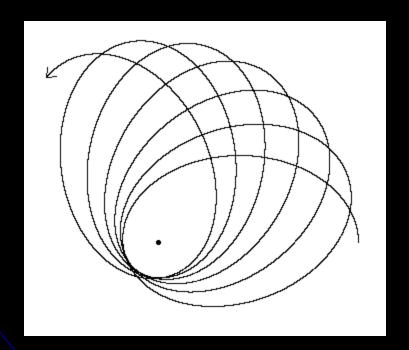


#### **Outline**

- I. Experimental Overview
- II. Newtonian Philosophy
- III. Gravitational Redshift
- IV. Ritzian Emission Theory
- V. Precession of Mercury
- VI. Bending of Starlight
- VII. Conclusions
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## Experimental Overview

1845 – Urbain LeVerrier observes a 35" per century excess precession of Mercury's orbit



Modern accepted value is 43.11" ± .45" per century Would take over 3,000,000 years to complete a circle

- 1915 Einstein's General Relativity predicts 42.9" precession
- 1919 Eddington measures deflection of starlight during total solar eclipse 1.75"
- 1960 Pound and Rebka measure a redshift of 2.5 x 10-15

## Newtonian Philosophy

- Absolute space and time
- Atomic Hypothesis
- Determinism
- Three laws of motion
- Action-at-a-Distance
- •"The whole burden of philosophy seems to consist in this- from the phenomena to investigate the forces of Nature and then from these forces to demonstrate the other phenomena"- Newton
- Unification of Force is implicit

#### Historical Interlude

- Gravitational Redshift
   1784 Rev. John Michell proposes using gravitational redshift to measure the mass of stars
- Bending of Starlight
   1785 Henry Cavendish calculates the deflection of light due to Sun's gravity, getting half the value of the modern value.
- Precession of Mercury
   1875 Tisserand uses gravitational version of Weber's law to arrive at correct value.

## Gravitational Redshift

$$E_{bottom} = E_{top} \left( 1 + \frac{gh}{c^2} \right)$$

$$1 + z = \frac{\lambda_{top}}{\lambda_{bottom}} = \frac{h v_{bottom}}{h v_{top}} = \frac{E_{bottom}}{E_{top}} = 1 + \frac{gh}{c^2}$$

From Pound & Rebka (1960)...

$$h = 22.5m \Rightarrow z = \frac{\Delta\lambda}{\lambda} = 2.5x10^{-15}$$

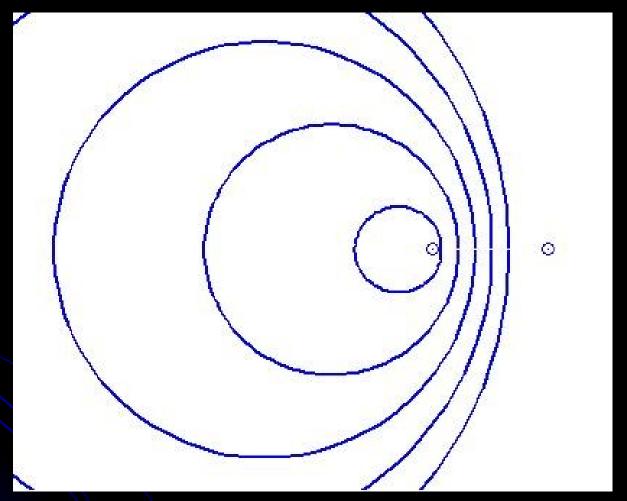
## Unification Scheme

- Gravity
- Thermodynamics
- Atomic Physics
- Electromagnetism

## Force Transmission

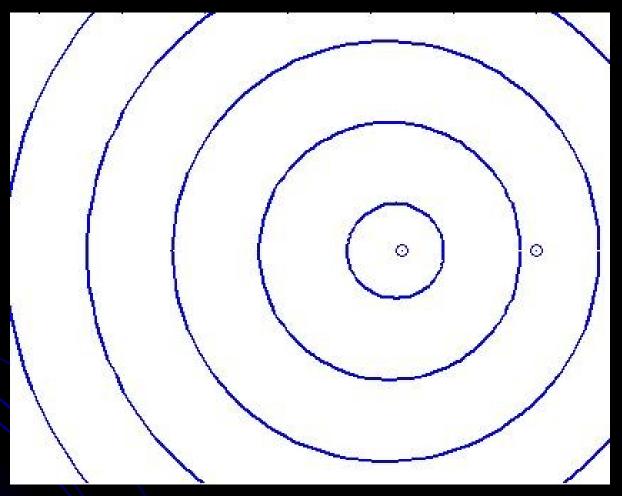
- Ballistic method
   Action-at-a-distance
- Medium methodEther

## Medium Picture



Waves move outward at velocity c, independent of source velocity

## Ballistic Picture



Particles move outward at velocity c+v'

## Ritzian Emission Theory



$$\vec{F}_{2\to 1} = \frac{Q_1 Q_2}{4\pi\varepsilon_0 r^2} \left\{ \left[ 1 + \frac{3-k}{4} \left( \frac{v}{c} \right)^2 - \frac{3(1-k)}{4c^2} (\vec{v} \cdot \hat{r})^2 - \frac{r}{2c^2} (\vec{a}_2 \cdot \hat{r}) \right] \hat{r} - \frac{k+1}{2c^2} (\vec{v} \cdot \hat{r}) \vec{v} - \frac{r}{c^2} \vec{a}_2 \right\}$$

#### Explains:

- Electrostatics
- Magnetostatics
- •Electrodynamics
- Radiation/Optics
- Radiation Reaction

## **Gravitational Version**

$$\frac{Q \to M}{\frac{1}{4\pi\varepsilon_0} \to G}$$

$$\vec{F}_{2\to 1} = G \frac{M_1 M_2}{r^2} \left\{ \left[ 1 + \frac{3 - k}{4} \left( \frac{v}{c} \right)^2 \dots \right] \right\}$$

#### Insert into Newton's 2nd Law

$$\vec{M_1 a_1} = G \frac{M_1 M_2}{r^2} \left\{ \left[ 1 + \frac{3 - k}{4} \left( \frac{v}{c} \right)^2 \dots \right] \right\}$$

## Precession of Mercury

$$\left| \frac{d^2x}{dt^2} = -\frac{\mu x}{r^3} \left\{ 1 + \frac{3 - k}{4c^2} \left[ \left( \frac{dx}{dt} \right)^2 + \left( \frac{dy}{dt} \right)^2 \right] - \frac{3(1 - k)}{4c^2} \left( \frac{dr}{dt} \right)^2 \right\} + \frac{\mu(k+1)}{2c^2 r^2} \frac{dx}{dt} \frac{dr}{dt}$$

$$\frac{d^2y}{dt^2} = -\frac{\mu y}{r^3} \left\{ 1 + \frac{3-k}{4c^2} \left[ \left( \frac{dx}{dt} \right)^2 + \left( \frac{dy}{dt} \right)^2 \right] - \frac{3(1-k)}{4c^2} \left( \frac{dr}{dt} \right)^2 \right\} + \frac{\mu(k+1)}{2c^2r^2} \frac{dy}{dt} \frac{dr}{dt}$$

#### Ugly... But not a tensor!

$$\left| \frac{d}{dt} \left( y \frac{dx}{dt} - x \frac{dy}{dt} \right) \right| = \frac{\mu(k+1)}{2c^2 r^2} \frac{dr}{dt} \left( y \frac{dx}{dt} - x \frac{dy}{dt} \right)$$

$$\left(y\frac{dx}{dt} - x\frac{dy}{dt}\right) = \alpha e^{\frac{-(k+1)\mu}{2c^2r}}$$

Angular Momentum

$$r^2 \frac{d\phi}{dt} = \alpha \left[ 1 - \frac{(k+1)\mu}{2c^2 r} \right]$$

## Mercury Cont'd

Add higher order terms to make equation of motion integrable

$$\mu \frac{d^2x}{dt^2} \left(\frac{k+1}{2c^2r}\right) + \mu x \frac{1-k}{2c^2r^2} \left(\frac{x}{r} \frac{d^2x}{dt^2} + \frac{y}{r} \frac{d^2y}{dt^2}\right) \\ \mu \frac{d^2y}{dt^2} \left(\frac{k+1}{2c^2r}\right) + \mu y \frac{1-k}{2c^2r^2} \left(\frac{x}{r} \frac{d^2x}{dt^2} + \frac{y}{r} \frac{d^2y}{dt^2}\right) \\ \mu \frac{d^2y}{dt^2} \left(\frac{k+1}{2c^2r}\right) + \mu y \frac{1-k}{2c^2r^2} \left(\frac{x}{r} \frac{d^2x}{dt^2} + \frac{y}{r} \frac{d^2y}{dt^2}\right) \\ \mu \frac{d^2y}{dt^2} \left(\frac{k+1}{2c^2r}\right) + \mu y \frac{1-k}{2c^2r^2} \left(\frac{x}{r} \frac{d^2x}{dt^2} + \frac{y}{r} \frac{d^2y}{dt^2}\right) \\ \mu \frac{d^2y}{dt^2} \left(\frac{k+1}{2c^2r}\right) + \mu y \frac{1-k}{2c^2r^2} \left(\frac{x}{r} \frac{d^2x}{dt^2} + \frac{y}{r} \frac{d^2y}{dt^2}\right) \\ \mu \frac{d^2y}{dt^2} \left(\frac{k+1}{2c^2r}\right) + \mu y \frac{1-k}{2c^2r^2} \left(\frac{x}{r} \frac{d^2x}{dt^2} + \frac{y}{r} \frac{d^2y}{dt^2}\right) \\ \mu \frac{d^2y}{dt^2} \left(\frac{k+1}{2c^2r}\right) + \mu y \frac{1-k}{2c^2r^2} \left(\frac{x}{r} \frac{d^2x}{dt^2} + \frac{y}{r} \frac{d^2y}{dt^2}\right) \\ \mu \frac{d^2y}{dt^2} \left(\frac{x}{r} \frac{d^2x}{dt^2} + \frac{y}{r} \frac{d^2x}{dt^2}\right) \\ \mu \frac{d^2x}{dt^2} \left(\frac{x}{r} \frac{d^2x}{dt^2} + \frac{y}{r} \frac{d^2x}{dt^2}\right$$

Multiply each by  $\frac{dx}{dt}$ ,  $\frac{dy}{dt}$ , add, integrate to find first integral of motion

$$\frac{1}{2} \left( \frac{dx}{dt} \right)^{2} + \frac{1}{2} \left( \frac{dy}{dt} \right)^{2} - \frac{\mu(k+1)}{2r} \left[ 1 - \frac{1}{2c^{2}} \left( \frac{dx}{dt} \right)^{2} - \frac{1}{2c^{2}} \left( \frac{dy}{dt} \right)^{2} \right] - \frac{\mu(1-k)}{2r} \left[ 1 - \frac{1}{2c^{2}} \left( \frac{dr}{dt} \right)^{2} \right] - \frac{\mu^{2}}{2c^{2}r^{2}} = \beta = \text{constant}$$

## Mercury Cont'd

$$\frac{\alpha^{2}}{r^{4}} \left(\frac{dr}{d\phi}\right)^{2} = -\frac{\alpha^{2}}{r^{2}} \left[ 1 + \frac{\mu(k-1)}{2c^{2}r} \right] + \frac{2\mu}{r} \left[ 1 + \frac{\mu(2k+1)}{2c^{2}r} \right] + 2\beta \left( 1 + \frac{\mu k}{c^{2}r} \right)$$

$$= \left[ 1 + \frac{\mu(k-1)}{2c^{2}r} \right] \times \left[ -\frac{\alpha^{2}}{r} + \frac{\mu^{2}(k+2)}{c^{2}r^{2}} + \frac{2\mu}{r} + \frac{\beta\mu(k+1)}{c^{2}r} + 2\beta \right]$$

let 
$$\frac{1}{r} = p$$
  $A = \frac{2\beta}{\alpha^2}$   $B = \frac{2\mu}{\alpha^2} + \frac{\beta\mu(k+1)}{\alpha^2c^2}$   $C^2 = 1 - \frac{\mu^2(k+2)}{\alpha^2c^2}$ 

$$\phi - \phi_0 = \int \frac{dp \left[ 1 - \frac{\mu(k-1)p}{4c^2} \right]}{\sqrt{-C^2p^2 + Bp + A}}$$

$$\phi - \phi_0 = \left[1 + \frac{\mu^2(k+5)}{4\alpha^2c^2}\right] \arcsin\left[\frac{2c^2p - B}{\sqrt{B^2 + 4AC^2}}\right] + \frac{\mu(k-1)}{4c^2}\sqrt{-C^2p^2 + Bp + A}$$

 $\phi$  advances not by  $\pi$  after half a revolution, but by  $\pi \left[1 + \frac{\mu^2(k+5)}{4\alpha^2c^2}\right]$ 

## Mercury Cont'd

If N is the number of revolutions per century, the angle of precession is

$$\delta = \frac{\mu^2 \pi (k+5) N}{2\alpha^2 c^2}$$

$$\frac{\mu}{\alpha^2} = \frac{1}{a(1 - e^2)}$$

 $a_0 \equiv$  average distance between Earth & Sun = 1 Au

$$v_0 \equiv \text{orbital speed} = 29.77 \frac{\text{km}}{\text{s}}$$

 $a \equiv$  average distance between Planet & Sun = .38703 Au  $e \equiv$  eccentricity of Planet = .206

$$N = 414.9378 \frac{\text{revs.}}{\text{century}}$$

$$\mu = a_0 v_0^2$$

$$\delta = \frac{\pi(k+5)}{2(1-e^2)} \left(\frac{v_0}{c}\right)^2 \frac{a_0}{a} N$$

$$\frac{2\delta(1-e^2)}{N\pi} \left(\frac{c}{v_0}\right)^2 \frac{a}{a_0} - 5 = k$$

## k = 7 results in a precession of 43.1" per century

...On to the Bending of Starlight...

## Bending of Starlight

$$\frac{d^2x}{dt^2} = -\frac{\mu x}{r^3} \left\{ 1 + \frac{3 - k}{4c^2} \left[ \left( \frac{dx}{dt} \right)^2 + \left( \frac{dy}{dt} \right)^2 \right] - \frac{3(1 - k)}{4c^2} \left( \frac{dr}{dt} \right)^2 \right\} + \frac{\mu(k+1)}{2c^2 r^2} \frac{dx}{dt} \frac{dr}{dt}$$

$$\frac{d^2y}{dt^2} = -\frac{\mu y}{r^3} \left\{ 1 + \frac{3-k}{4c^2} \left[ \left( \frac{dx}{dt} \right)^2 + \left( \frac{dy}{dt} \right)^2 \right] - \frac{3(1-k)}{4c^2} \left( \frac{dr}{dt} \right)^2 \right\} + \frac{\mu(k+1)}{2c^2r^2} \frac{dy}{dt} \frac{dr}{dt}$$

Add higher order terms to x-component to make equation integrable

$$\frac{\mu}{c^2r} \left( \ddot{x} - \frac{\mu x}{r^3} \right) + \frac{\mu x}{c^2r^3} \left( x\ddot{x} + y\ddot{y} - \frac{\mu}{r} \right)$$
 Similar term for y-component

Energy Conservation Equation

$$\frac{1}{2}v^{2} + \frac{(1+k)\mu}{2r} \left(1 - \frac{v^{2}}{2c^{2}}\right) + \frac{(1-k)\mu}{2r} \left(1 - \frac{r^{2}}{2c^{2}}\right) - \mu^{2}/2c^{2}r^{2} = \text{constant} = \frac{1}{2}\omega^{2}$$

## Bending of Starlight

Angular Momentum Conservation Equation

$$J = h \left( 1 + \frac{1+k}{2} \frac{\mu \rho}{c^2} \right)$$

where 
$$h = \omega b$$
,  $\rho = \frac{1}{r}$ ,  $b \equiv \text{impact parameter}$ 

Using b = 1 solar radius  $(6.96x10^8 \text{ m})$  and k=7

Angular Deflection Equation

$$\Delta \theta = \frac{4GM_S}{bc^2} = 1.75$$
"

## Conclusions

- Explanations of the precession of Mercury, the bending of starlight and gravitational redshift do <u>not</u> require the General Theory of Relativity.
- Ritz's Theory explains both observations if the constant k has the value 7.

### **Future Work**

- Gravitational Lensing
- Gravitational Waves
- Unify Gravity and Electromagnetism

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#### G.R. Parameters vs. Our Parameters



**(=7**